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# HALL THRUSTERS INSTABILITY: NUMERICAL STUDY

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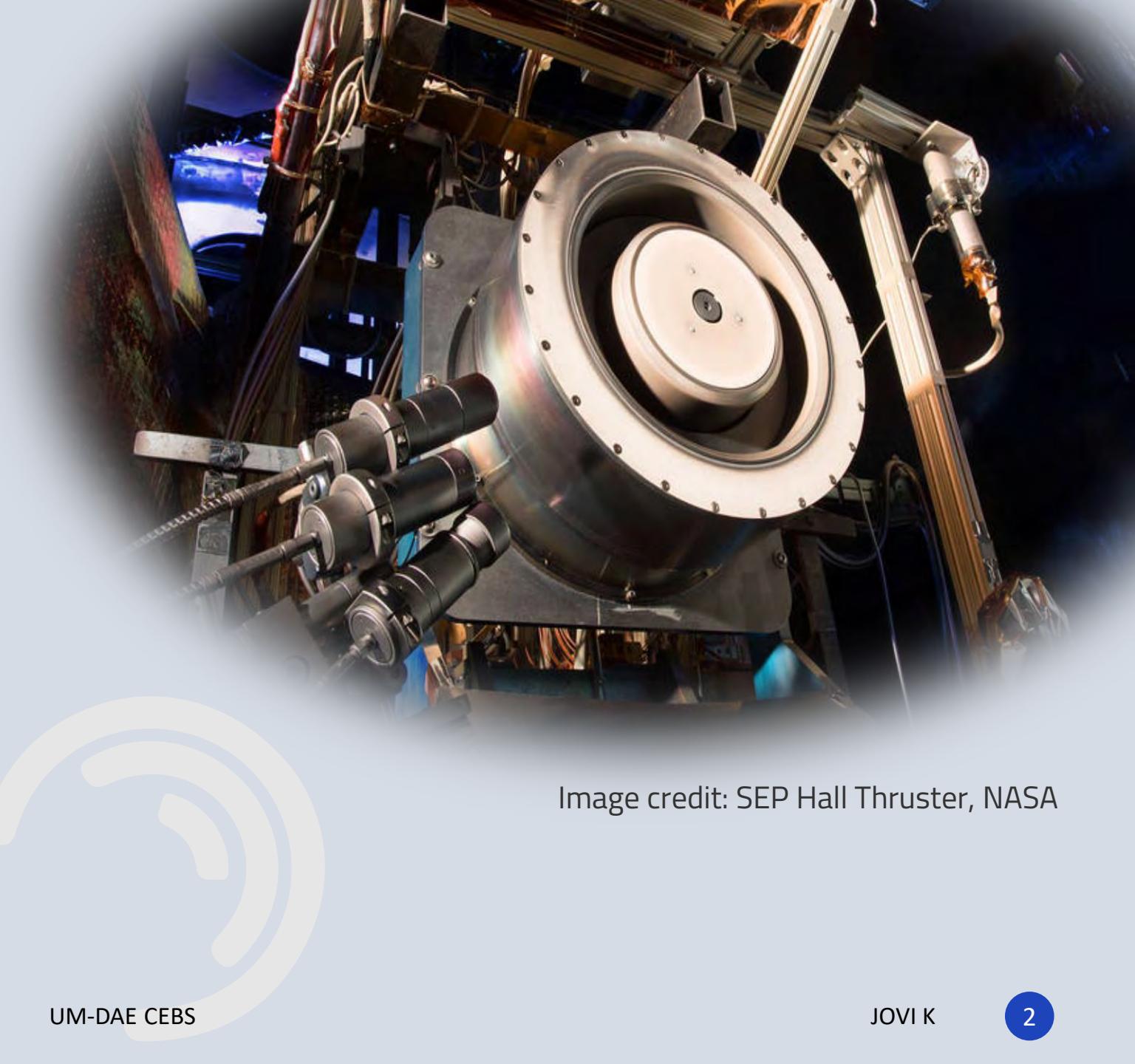


Image credit: SEP Hall Thruster, NASA

# SESSION 1

## INTRODUCTION

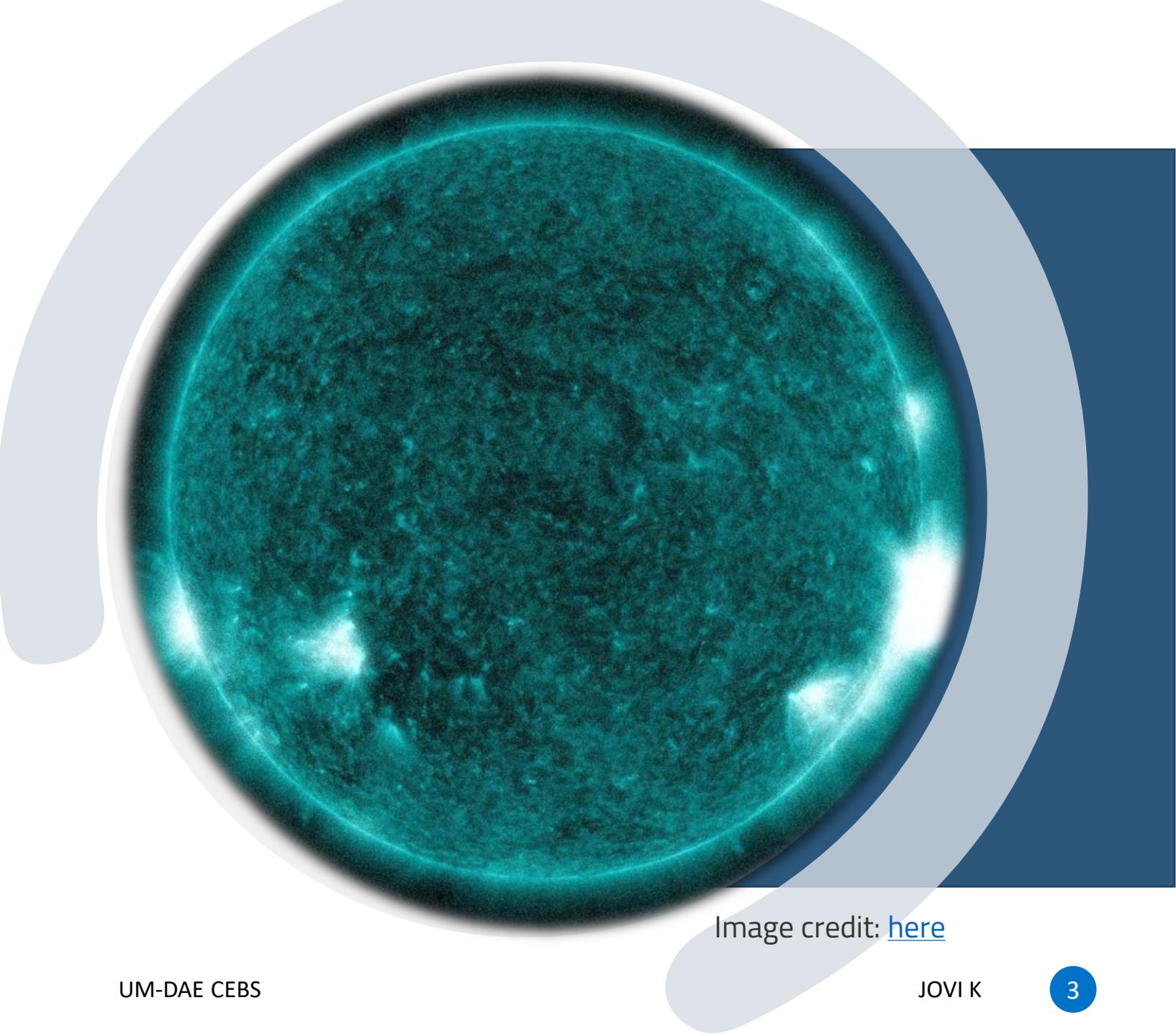


Image credit: [here](#)

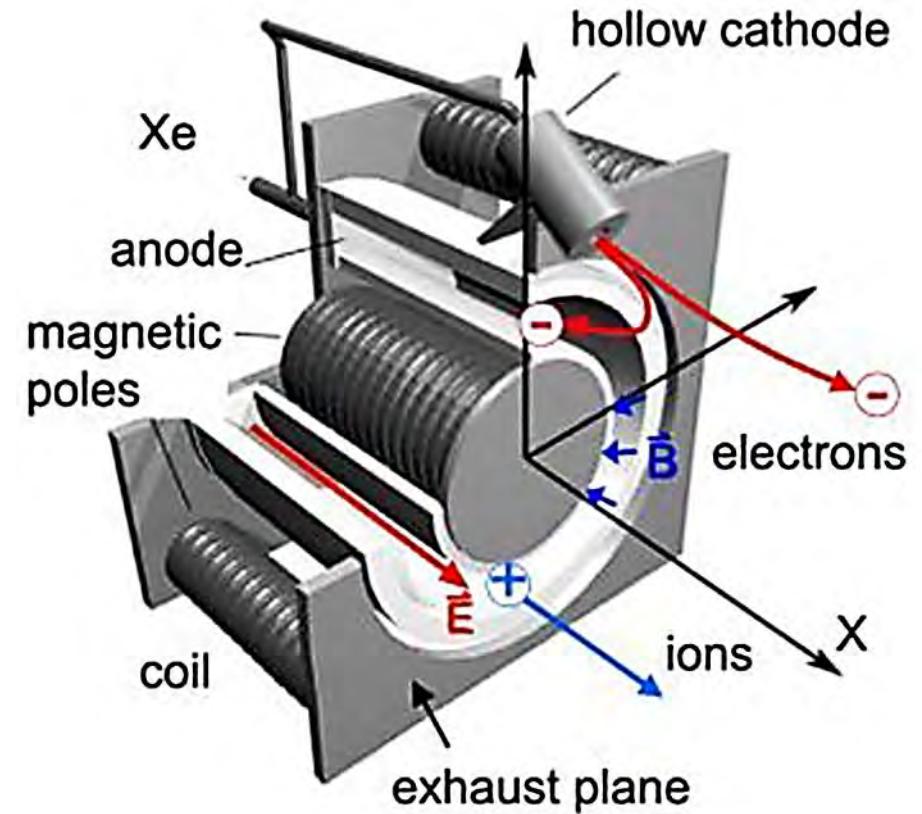
# HALL THRUSTERS

## PRINCIPLE

Gridless Ion Thrusters that makes use of **Hall effect**.

Permanent magnets provide the external magnetic barrier for electrons.

Confined electrons ionizes the neutral atoms



## SCHEMATIC OF HALL THRUSTER

Image credit: L. Dubois, et. al.; 2018



## THE HALL THRUSTER PROBLEM

# MOTIVATION

The electron transport across the magnetic confinement  
**(ANOMALOUS TRANSPORT)** is a mystery.

Anything ranging from Electron collisions with the walls and secondary electron emission to instabilities and turbulence could be responsible for this anomalous transport along the magnetic field.

# SESSION 2

## DISPERSION RELATION

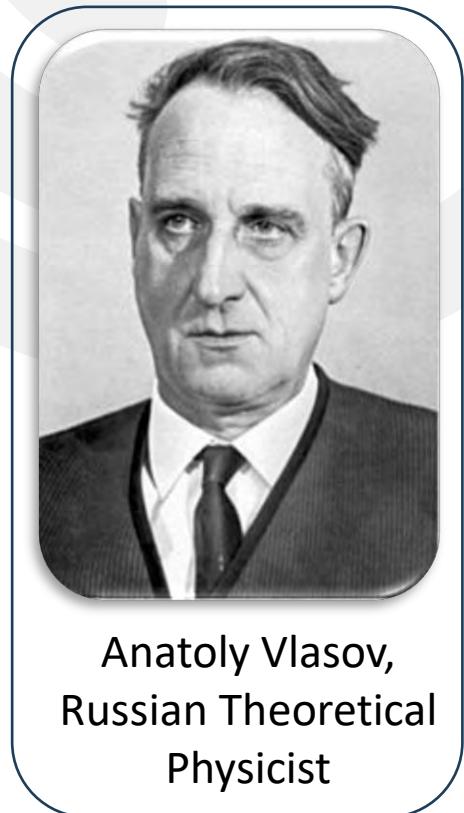


# VLASOV - MAXWELL EQUATIONS

Time evolution of the distribution function.

Collisionless Vlasov - Maxwell equation,

$$\frac{\partial f_\sigma}{\partial t} + \nabla \cdot \mathbf{v}_\sigma f_\sigma - q_\sigma \frac{\partial}{\partial \mathbf{p}} \cdot \left( \mathbf{E} + \frac{\mathbf{v}_\sigma}{c} \times \mathbf{B} \right) f_\sigma = 0$$



Anatoly Vlasov,  
Russian Theoretical  
Physicist

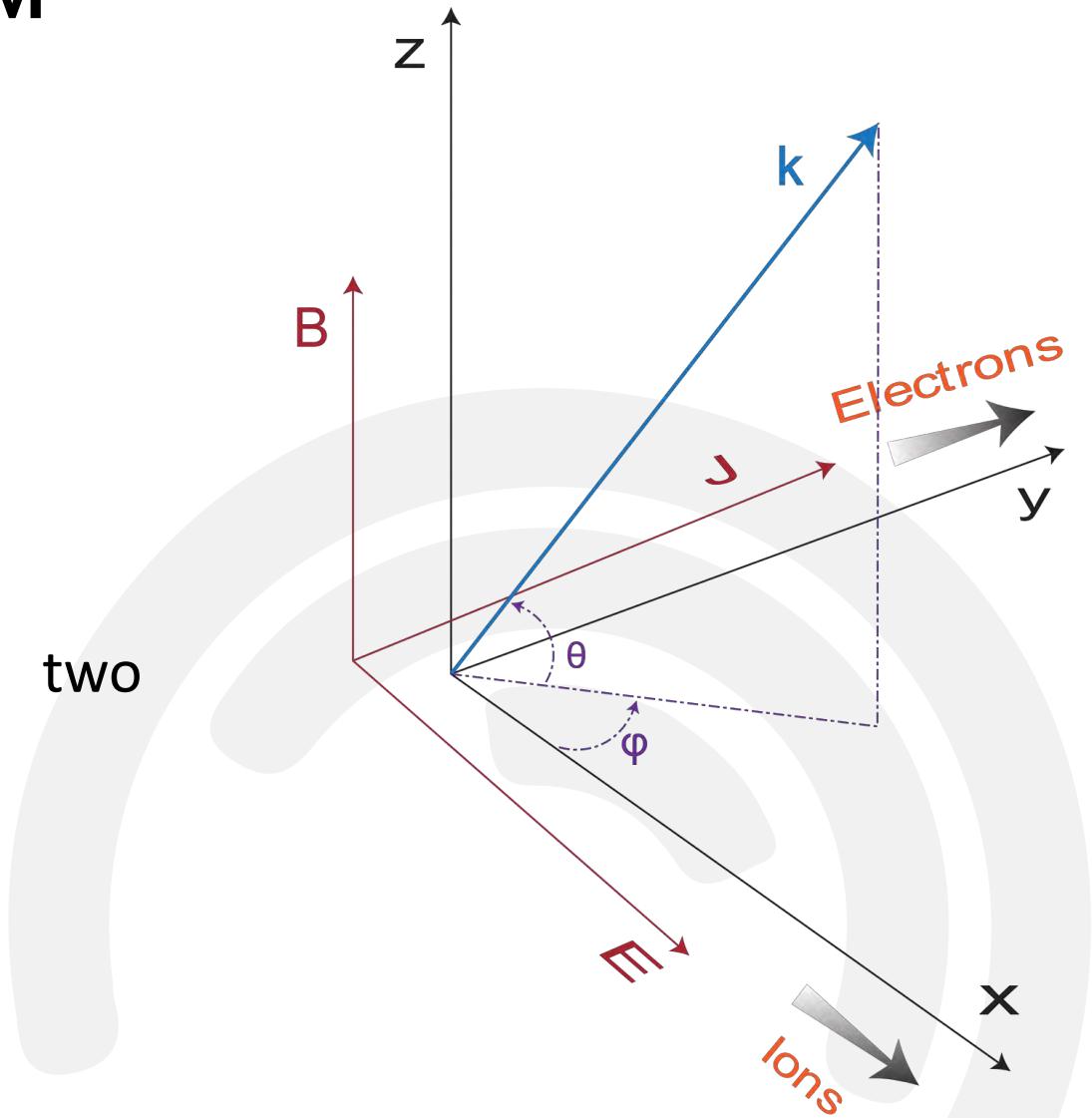
# THE HALL THRUSTER PROBLEM

## Linear Perturbation Model

- Localized Cartesian Coordinates.

$$\omega_{ce} \gg |\omega| \gg \omega_{ci}$$

- separation of charges along the two perpendicular directions



# SCHEMATICS

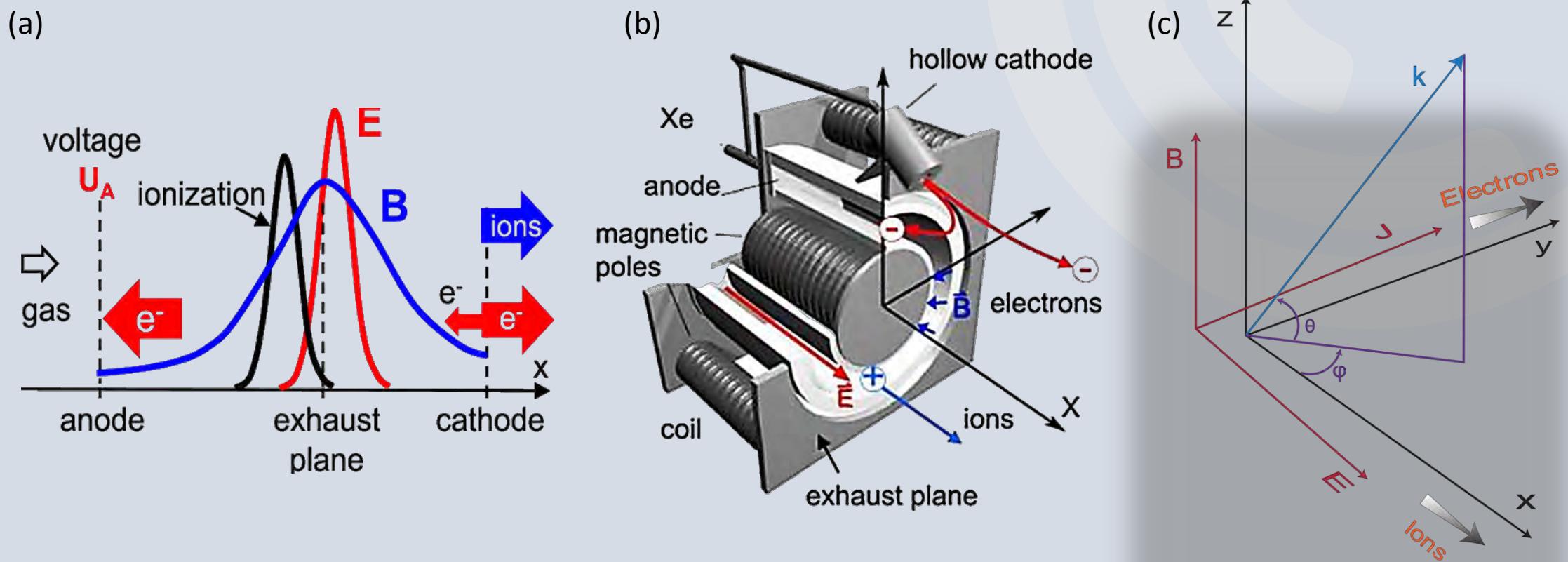


FIG 1: (a) Intensity plot with axial distance; (b) schematic of the Hall exhaust and; (c) localized cartesian coordinates

# DISPERSION RELATION

**Warm magnetized plasma (Doppler shifted electrons) electrostatic dispersion relation ,**

$$\frac{1}{T_e} \left[ 1 + \frac{k^2 v_{Te}^2}{\omega_{pe}^2} \right] + \frac{1}{T_i} \left[ 1 + \frac{\omega}{\sqrt{2} k_z v_{Ti}} Z \left( \frac{\omega}{\sqrt{2} k_z v_{Ti}} \right) \right] + \frac{(\omega - k_y v_0)}{\sqrt{2} k_z v_{Te}} \sum_{n=-\infty}^{\infty} e^{-k_\perp^2 r_{L\sigma}^2} I_n(k_\perp^2 r_{L\sigma}^2) Z \left( \frac{\omega - k_y v_0 - n\omega_{ce}}{\sqrt{2} k_z v_{Te}} \right) = 0$$

$I_n$  is Modified Bessel's Function of first kind

Plasma Distribution Function,

$$Z(\alpha) = \frac{1}{\sqrt{\pi}} \int_{-\infty}^{\infty} d\xi \frac{\exp(-\xi^2)}{(\xi - \alpha)}$$

# NORMALIZED DISPERSION RELATION

Warm magnetized plasma (Doppler shifted electrons) electrostatic dispersion relation ,

$$\hat{k}^2 + \frac{1}{\hat{T}} \left[ 1 + \frac{\Omega_i}{\sqrt{2 Y_i}} Z \left( \frac{\Omega_i}{\sqrt{2 Y_i}} \right) \right] + [1 + g(\Omega_e, X_e, Y_e)] = 0$$

$$\Omega_e = (\hat{\omega} - \hat{k}_y \hat{v}_d) / \hat{\omega}_{ce}$$
$$\Omega_i = (\hat{\omega} - \hat{k}_x \hat{v}_p) / \hat{\omega}_{ci}$$
$$; \quad X_e = \frac{\hat{k}_\perp^2}{\hat{\omega}_{ce}^2} \hat{M}$$

$$Y_i = \hat{k}_z^2 \hat{T}$$
$$Y_e = \frac{\hat{k}_z^2}{\hat{\omega}_{ce}^2} \hat{M}$$

$$\hat{T} = \frac{T_i}{T_e}; \quad \hat{M} = \frac{m_i}{m_e}$$

# SESSION 3

## Numerical Solutions

# NUMERICAL ANALYSIS

## PARAMETERS

$M_i$ (kg)	$E_0$ (V/m)	$B_0$ (T)	$n_e$ ( $m^{-3}$ )	$T_e$ (eV)	$v_p$ (m/s)	$T_i$ (eV)
$2.2 \times 10^{-25}$	$1 \times 10^4$	$15 \times 10^{-3}$	$2 \times 10^{17}$	25	16,000	10

Table 1: Typical parameters at the exit plane of the Snecma 5kW PPSX000VR Hall thruster.

# NUMERICAL ANALYSIS

## 1. DUCROCQ EQUATION

1. DUCROCQ EQUATION
2. MODIFIED DUCROCQ EQUATION
3. GENERAL DISPERSION RELATION

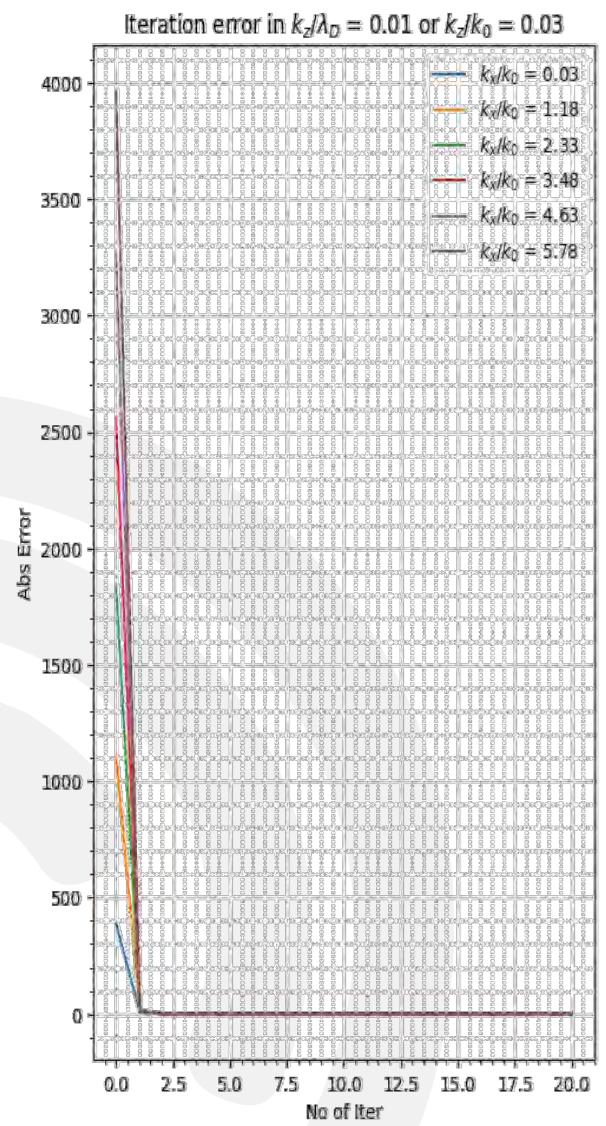
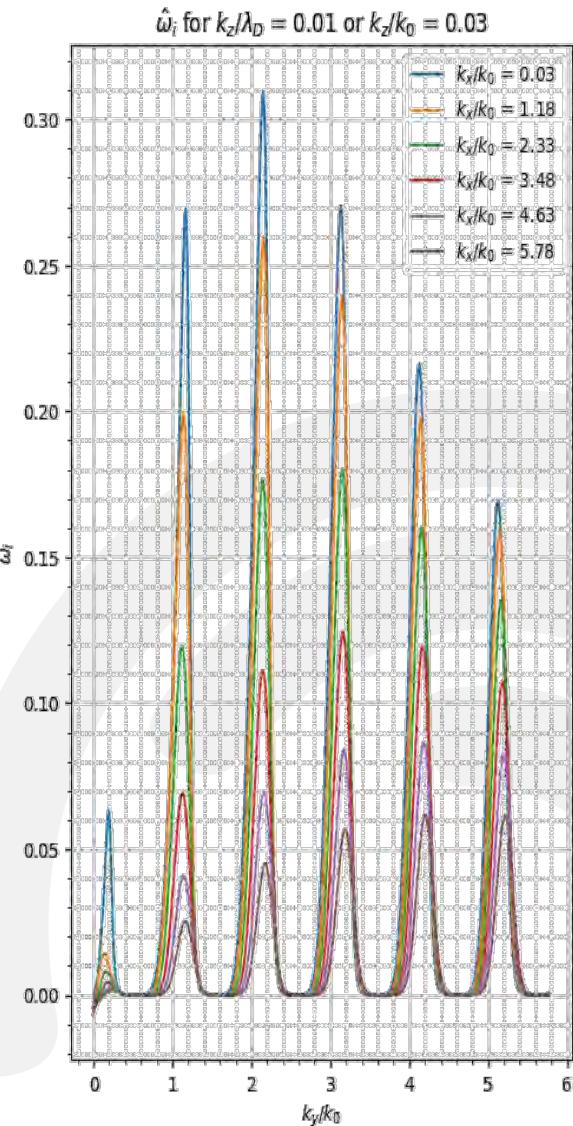
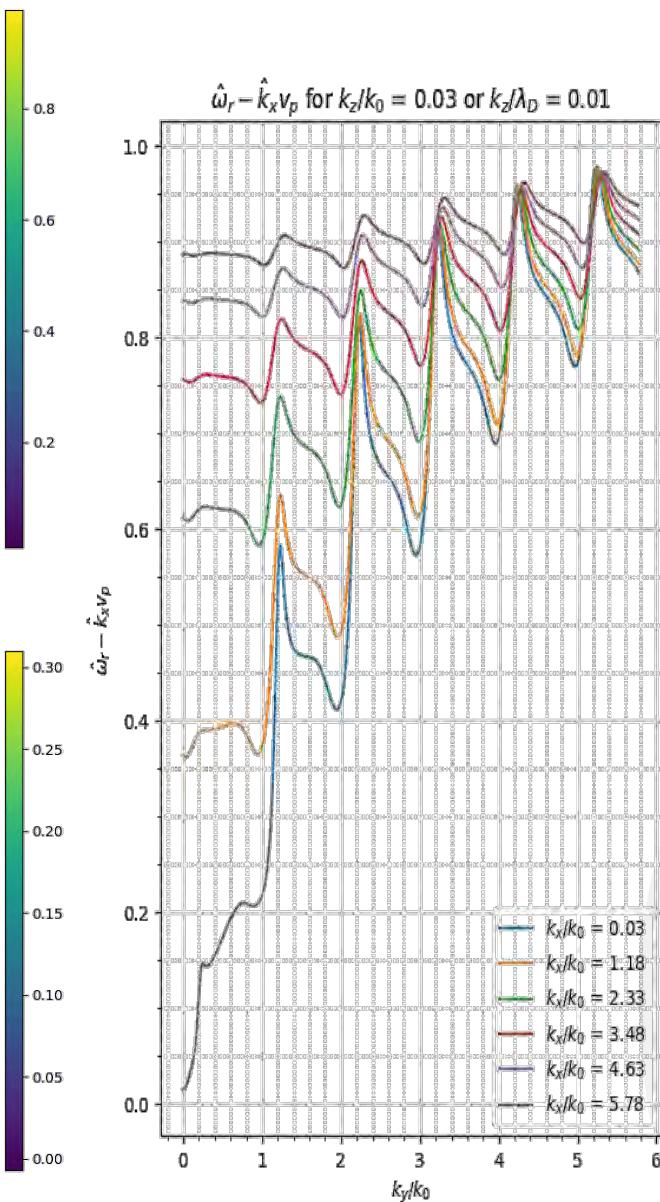
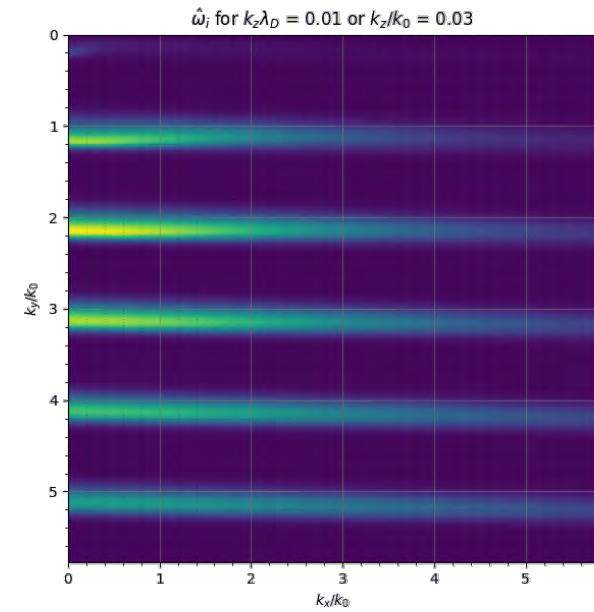
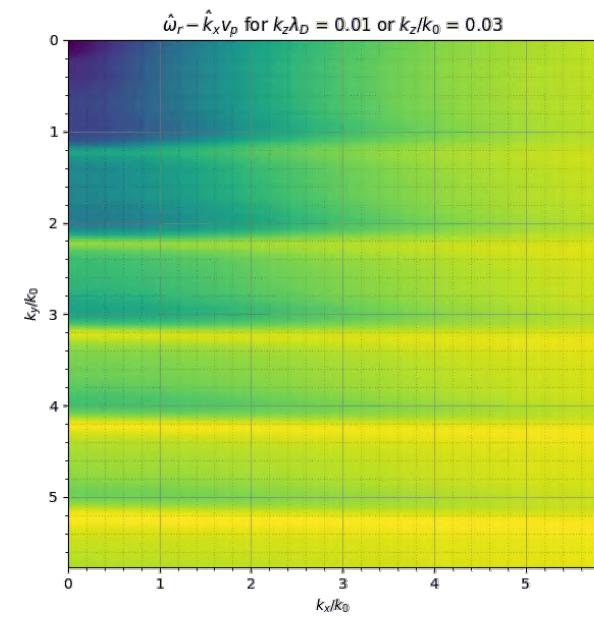
For  $\hat{T} \ll 1$ , the general dispersion becomes (to 0<sup>th</sup> order in  $\hat{T}$ ),

$$1 + k^2 \lambda_D^2 + g\left(\frac{\omega - k_y V_d}{\omega_{ce}}, (k_x^2 + k_y^2)\rho^2, k_z^2\rho^2\right) - \frac{k^2 \lambda_D^2 \omega_{pi}^2}{(\omega - k_x v_p)^2} = 0$$

Normalised iterative form,

$$(\hat{\omega}_{n+1} - \hat{k}_x \hat{v}_p)^2 = \hat{k}^2 \Bigg/ 1 + \hat{k}^2 + g\left(\frac{\hat{\omega}_n - \hat{k}_y \hat{V}_d}{\hat{\omega}_{ce}}, (\hat{k}_x^2 + \hat{k}_y^2) \frac{\hat{M}}{\hat{\omega}_{ce}^2}, \hat{k}_z^2 \frac{\hat{M}}{\hat{\omega}_{ce}^2}\right)$$

# DUCROCQ EQUATION - $k_z$ dependance



# NUMERICAL ANALYSIS

## 2. MODIFIED DUCROCQ EQUATION

1. DUCROCQ EQUATION
2. MODIFIED DUCROCQ EQUATION
3. GENERAL DISPERSION RELATION

For  $\hat{T} \ll 1$ , the general dispersion becomes (to 1<sup>th</sup> order in  $\hat{T}$ ),

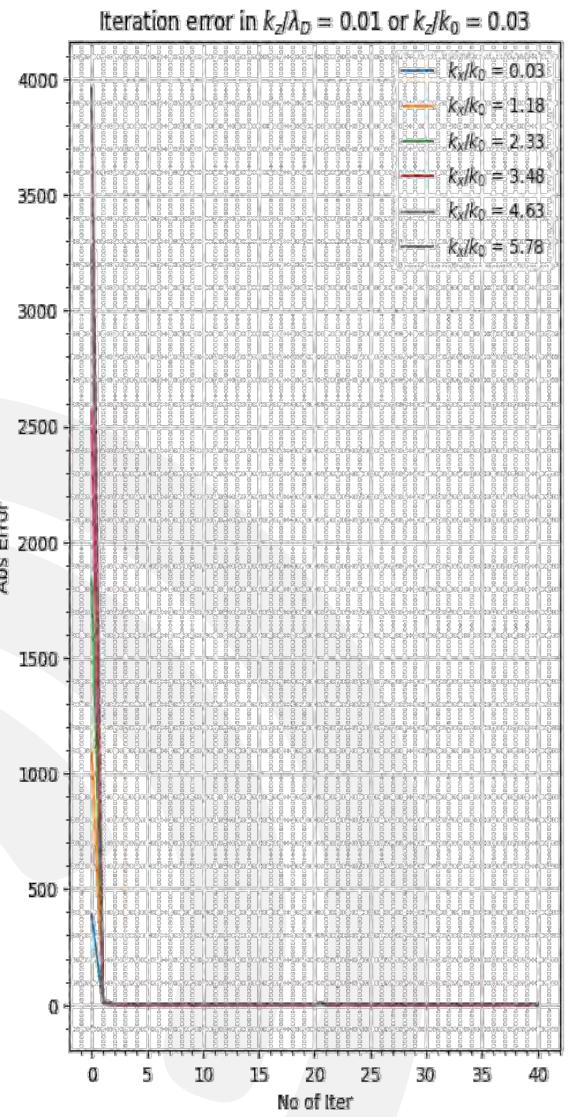
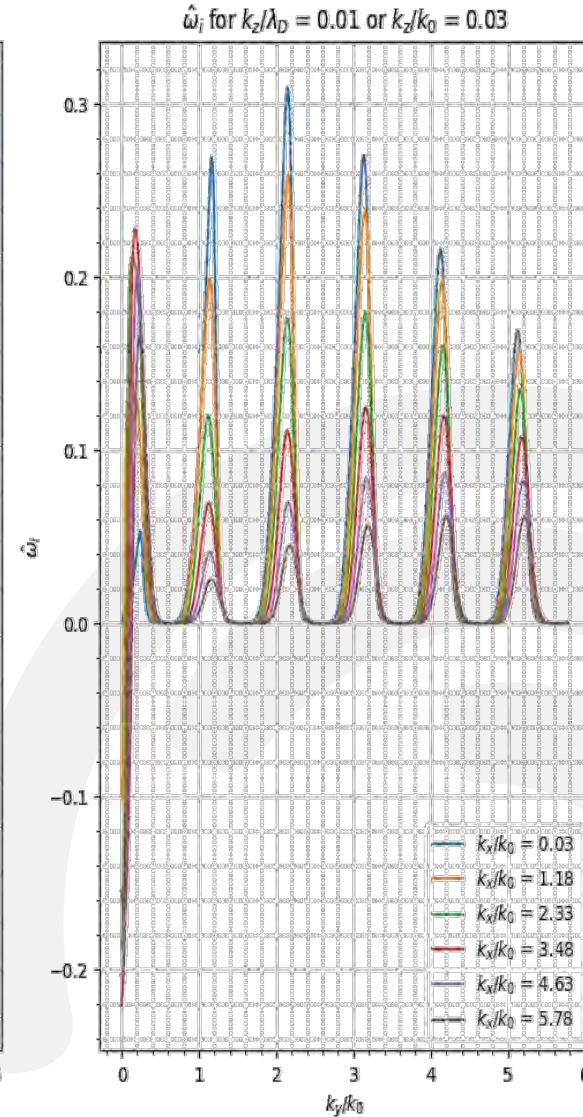
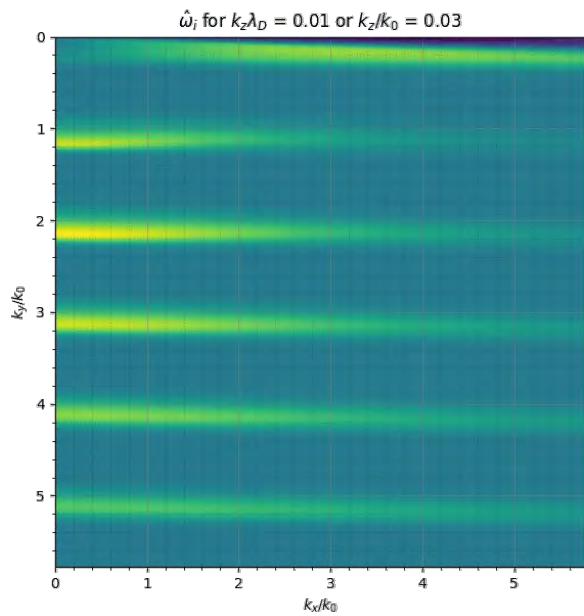
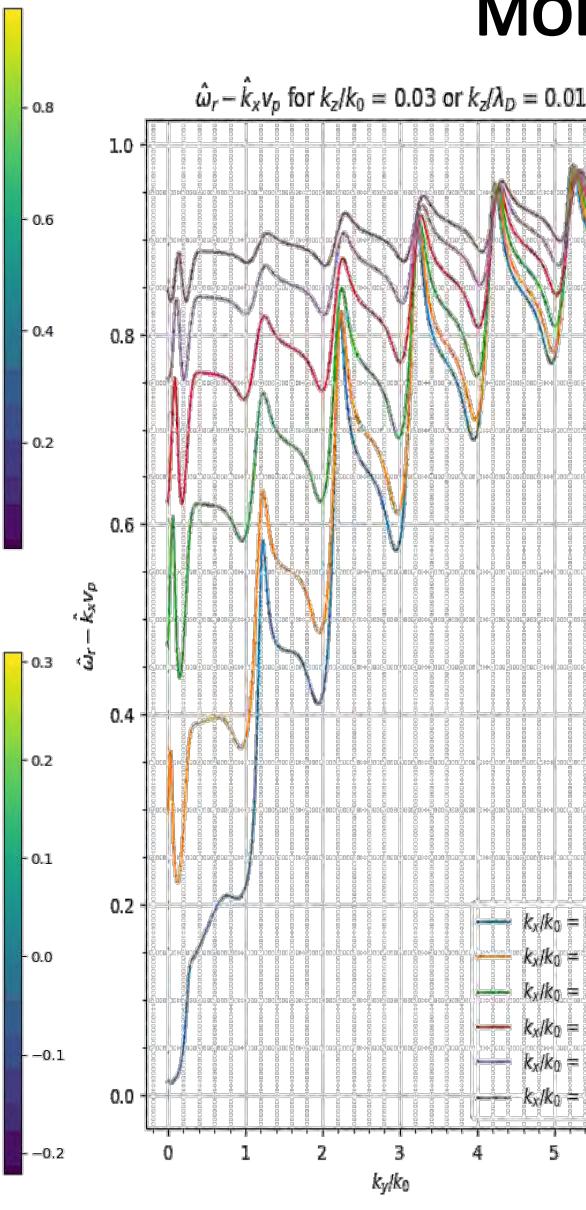
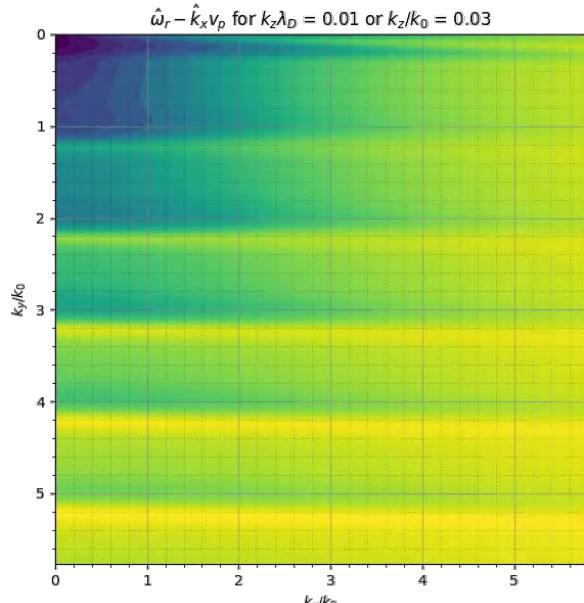
$$1 + k^2 \lambda_{De}^2 + g\left(\frac{\omega - k_y V_d}{\omega_{ce}}, (k_x^2 + k_y^2)\rho^2, k_z^2\rho^2\right) - \frac{T_e}{T_i} \left[ \frac{k^2 \lambda_{Di}^2 \omega_{pi}^2}{(\omega - k_x v_p)^2} - X \right] = 0$$

were  $X = \frac{i\sqrt{\pi}(\omega - k_x v_p)\sigma}{\sqrt{2}k_z v_{Ti}} \exp - \left( \frac{\omega - k_x v_p}{\sqrt{2}k_z v_{Ti}} \right)^2$

Normalised iterative form,

$$\begin{aligned} & (\hat{\omega}_{n+1} - \hat{k}_x \hat{v}_p)^2 \\ &= \hat{k}^2 \Bigg/ 1 + \hat{k}^2 + g\left(\frac{\hat{\omega}_n - \hat{k}_y \hat{V}_d}{\hat{\omega}_{ce}}, (\hat{k}_x^2 + \hat{k}_y^2) \frac{\hat{M}}{\hat{\omega}_{ce}^2}, \hat{k}_z^2 \frac{\hat{M}}{\hat{\omega}_{ce}^2}\right) + \frac{i\sqrt{\pi}\sigma}{\hat{T}} \frac{\Omega_{i,n}}{\sqrt{2} Y_i} \exp - \left( \frac{\Omega_{i,n}}{\sqrt{2} Y_i} \right)^2 \end{aligned}$$

# MODIFIED DUCROCQ EQUATION - $k_z$ dependance



# NUMERICAL ANALYSIS

## 3. GENERAL DISPERSION RELATION

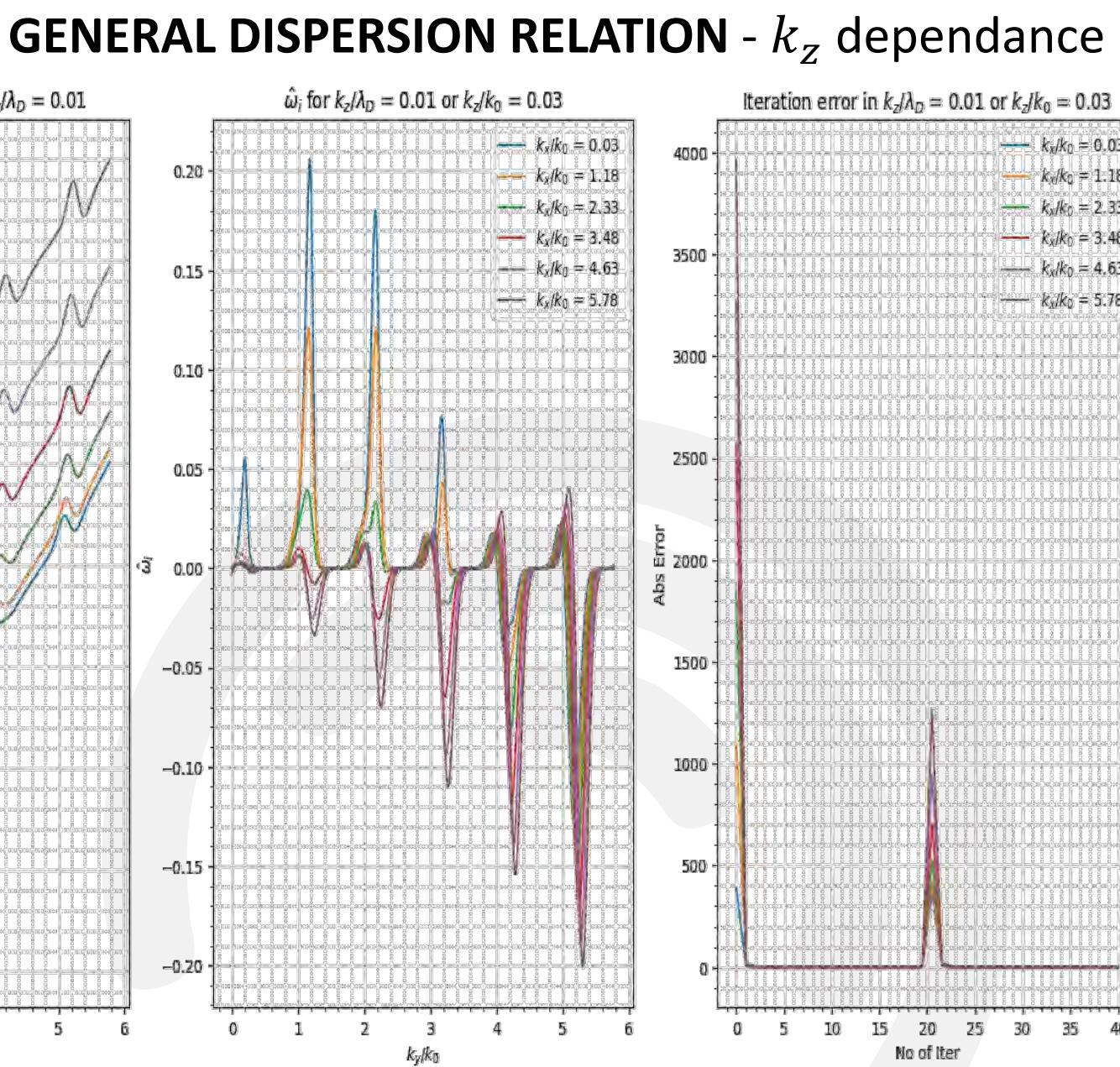
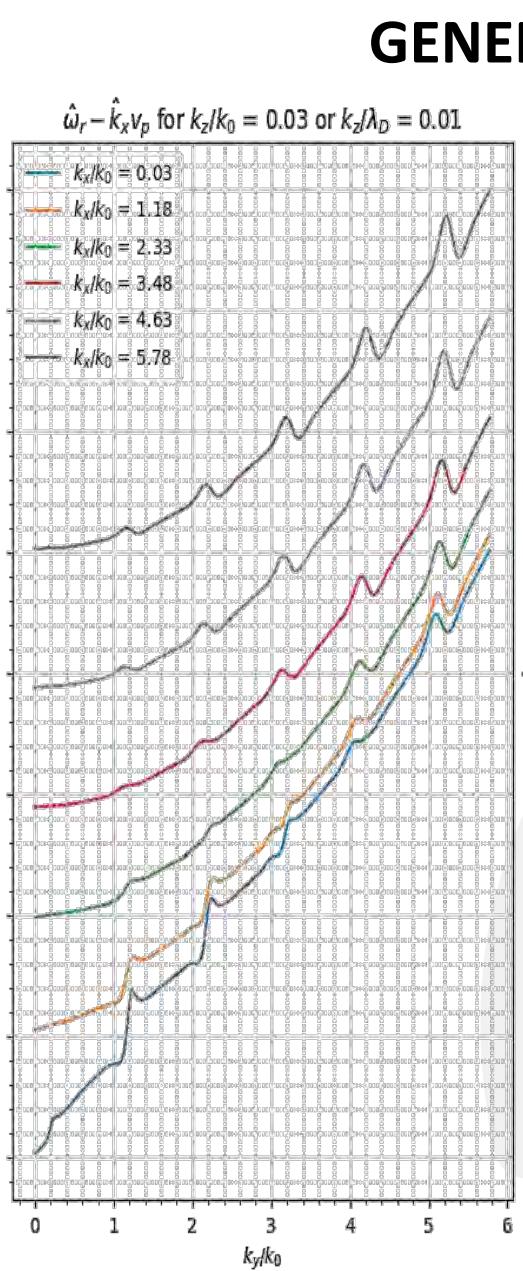
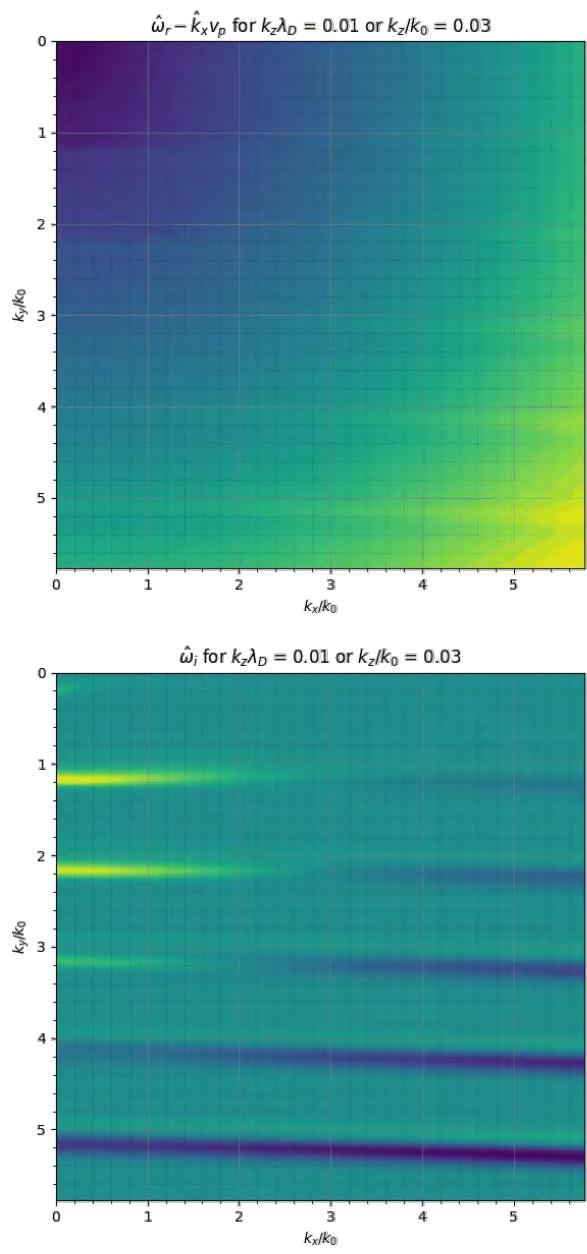
1. DUCROCQ EQUATION
2. MODIFIED DUCROCQ EQUATION
3. GENERAL DISPERSION RELATION

We have,

$$\hat{k}^2 + \frac{1}{\hat{T}} \left[ 1 + \frac{\Omega_i}{\sqrt{2 Y_i}} Z \left( \frac{\Omega_i}{\sqrt{2 Y_i}} \right) \right] + [1 + g(\Omega_e, X_e, Y_e)] = 0$$

Normalised iterative form,

$$\hat{\omega}_{n+1} = - \frac{\hat{\omega}_{ci} \sqrt{2 Y_i} \left\{ \hat{T} \left( 1 + \hat{k}^2 + g(\Omega_{n,e}, X_e, Y_e) \right) + 1 \right\}}{Z \left( \Omega_{n,i} / \sqrt{2 Y_i} \right)} + \hat{k}_x \hat{v}_p$$



# FUTURE PROSPECTS

- Temperature dependant analysis.
  - More vigorous treatment for the Numerical stability of the General dispersion relation.
  - Effect of perturbation on the initial velocity distributions.
  - Higher order Approximations to include highly non linear behaviour.
- 
- **CAN THESE EXPLAIN THE ANOMOLOUS ELECTRON TRANSPORT PROBLEM?**

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Image credit: [here](#)  
UM DAE CEBS



# THANK YOU

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SEM 10 Project Presentation

THE END



# APPENDIX



Instabilities



Plasma Disp.  
Function

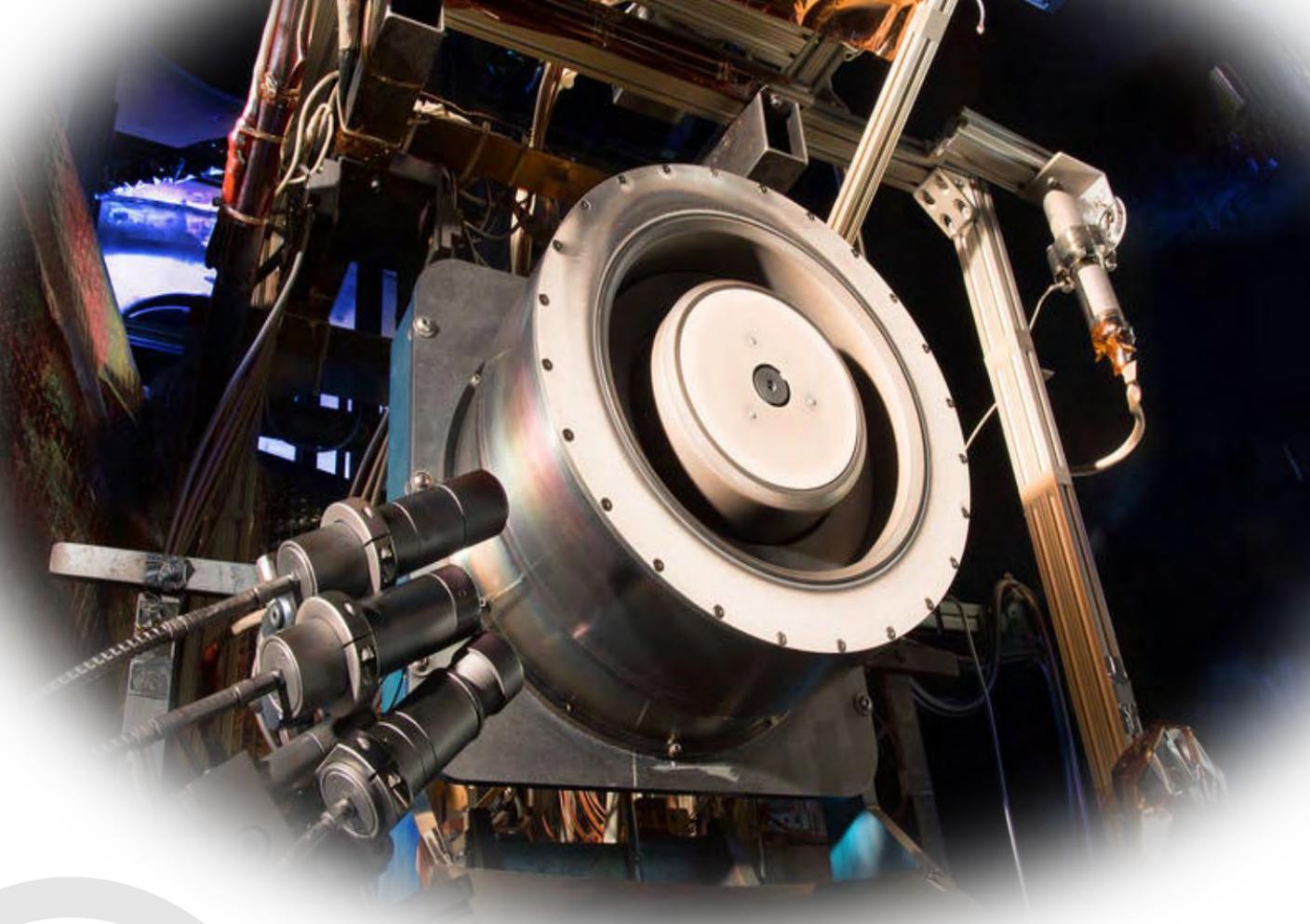


Image credit: SEP Hall Thruster, NASA

# OUTLINE OF DERIVATION

## Perturbed Vlasov Equation:

Under the Electrostatic Perturbations,

$$\begin{aligned}\vec{B} &= \mathbf{B} + 0 \\ \vec{E} &= \mathbf{E} - \nabla\phi_1 \\ f_\sigma &= f_{\sigma 0} + f_{\sigma 1}\end{aligned}$$

Perturbation is of the form,  
 $x_1 \sim \exp i(\mathbf{k} \cdot \mathbf{x} - \omega t)$

where  $\mathbf{B}$  and  $\mathbf{E}$  are initial uniform fields, the linearized Vlasov equation takes the form,

$$\frac{\partial f_{\sigma 1}}{\partial t} + \frac{\partial}{\partial \mathbf{x}} \cdot \mathbf{v} f_{\sigma 1} + \frac{q_\sigma}{m_\sigma} \frac{\partial}{\partial \mathbf{v}} \cdot (\mathbf{v} \times \mathbf{B} + \mathbf{E}) f_{\sigma 1} = \frac{q_\sigma}{m_\sigma} \nabla\phi_1 \cdot \frac{\partial f_{\sigma 0}}{\partial \mathbf{v}}$$

# ANALYTICAL INSTABILITY ANALYSIS

## COLD PLASMA LIMIT –

For the limit  $|(\omega - k_y v_0 - n\omega_{c\sigma})/k_{\parallel}v_T| \gg 1$ ,

$$1 - \frac{k_z^2}{k^2} \frac{\omega_{pi}^2}{\omega^2} - \frac{e^{-k_{\perp}^2 r_{Le}^2}}{k^2 \lambda_{De}^2} \left[ - \sum_{n=1}^{\infty} \frac{2n_e^2 \omega_{ce}^2}{(\omega - k_y v_0)^2 - n_e^2 \omega_{ce}^2} I_n(k_{\perp}^2 r_{Le}^2) \right] = 0$$

# ANALYTICAL INSTABILITY ANALYSIS

## COLD PLASMA LIMIT

### Modified Two Stream Instability –

For the case where  $\omega - k_y v_0 \ll \omega_{ce}$  (ions are unmagnetized),

To 2<sup>nd</sup> order we have,

$$1 - \frac{\omega_{pi}^2}{\omega^2} - \frac{\omega_{pe}^2 k_z^2}{(\omega - k_y v_0)^2 k^2} + \frac{\omega_{pe}^2 k_y^2}{\omega_{ce}^2 k^2} = 0$$

Coupling between  $\omega^2 \sim \omega_{pi}^2$ , and the doppler shifted electron modes,

$$(\omega - k_y v_0)^2 \sim \frac{\omega_{pe}^2 k_z^2}{k_y^2 (\omega_{pe}^2 / \omega_{ce}^2) + k^2}$$

# ANALYTICAL INSTABILITY ANALYSIS

## COLD PLASMA LIMIT

Modified Two Stream Instability

**Electron Cyclotron Drift Instability –**

For  $k_{\parallel}/k_{\perp} \ll 1$ ,

To 2<sup>nd</sup> order we have,

$$k_y^2 \left( 1 - \frac{\omega_{pe}^2}{(\omega - k_y v_{e0})^2 - \omega_{ce}^2} - \frac{\omega_{pi}^2}{\omega^2} \right) = 0$$

Coupling between  $\omega^2 = \omega_{pe}^2$ , and the doppler shifted coupling of electrons,

$$(\omega - k v_0)^2 = \omega_{pe}^2 + \omega_{ce}^2$$

# ANALYTICAL INSTABILITY ANALYSIS

## COLD PLASMA LIMIT

Modified Two Stream Instability

**Electron Cyclotron Drift Instability –**

To 2<sup>nd</sup> order we have,

$$k_y^2 \left( 1 - \frac{\omega_{pe}^2}{(\omega - k_y v_{e0})^2 - \omega_{ce}^2} - \frac{\omega_{pi}^2}{\omega^2} \right) = 0$$

This may be thought of as coupling between  $\omega^2 = \omega_{pe}^2$ , and the doppler shifted coupling of electrons,

$$(\omega - k v_0)^2 = \omega_{pe}^2 + \omega_{ce}^2$$